

ACCRETION AND OUTFLOW ACTIVITY
IN YOUNG CLUSTERS
AND
THE PROBLEM OF SKY SUBTRACTION

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Giusi Micela INAF - OAPa
... (project on wiki)

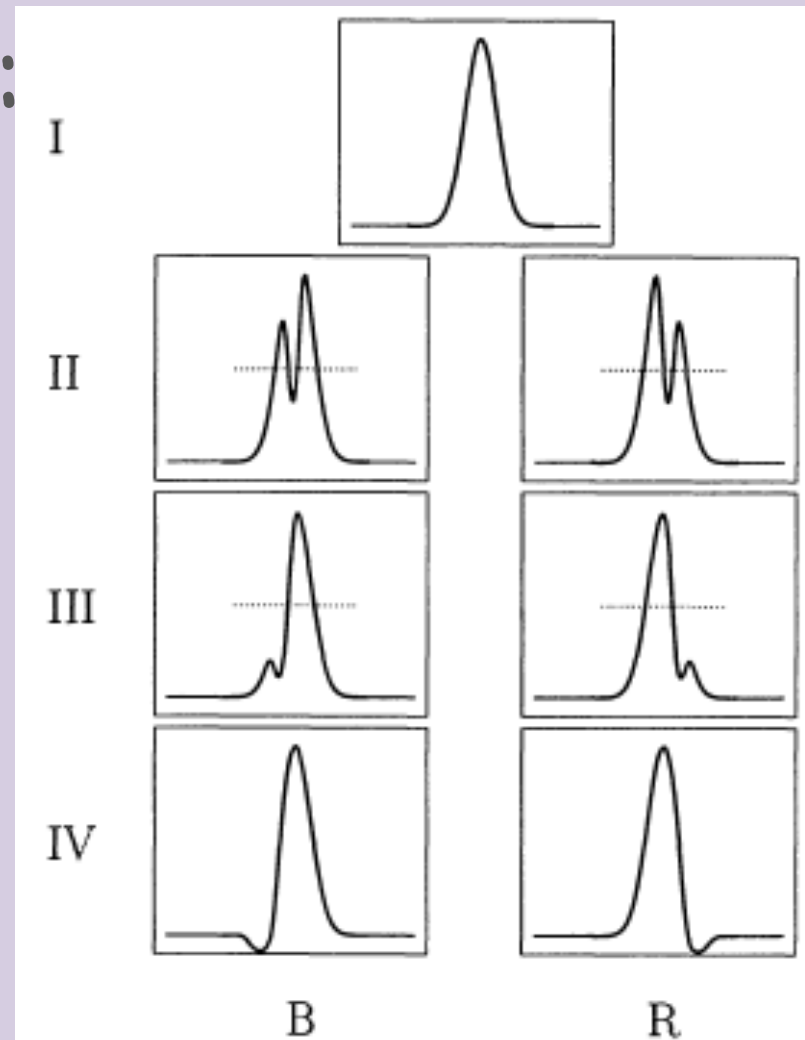
YOUNG CLUSTERS IN GES

- ◆ Study of accretion/outflow activity in young cluster members (H α and forbidden emission lines, FELs)
- ◆ Method for
Objects with H α emission and strong Nebular contribution to Accretion/outflow activity (OHaNA)
(several GES clusters:
NGC 2264, NGC 6530 - Prisinzano et al. 2007,
NGC 6611 - Bonito et al. 2013, ...)

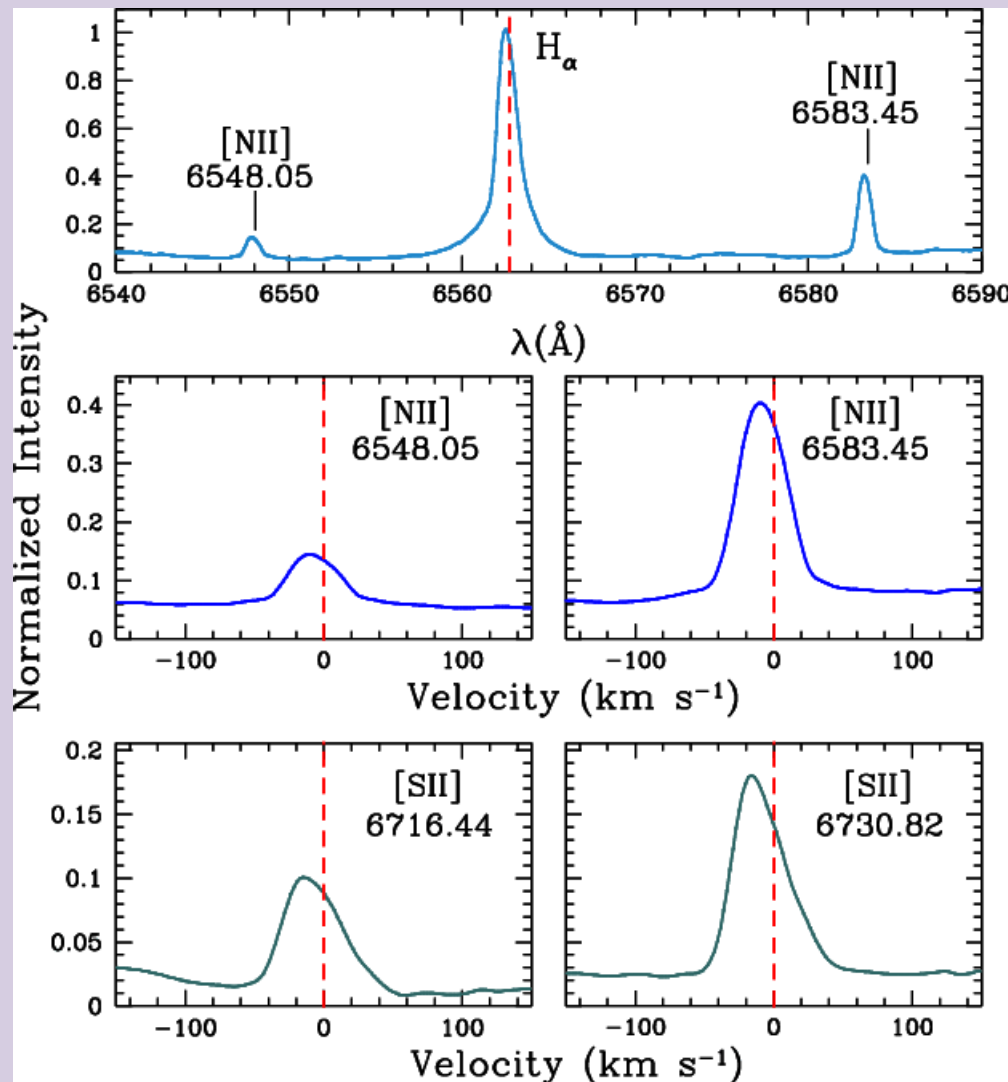
H α EMISSION LINE AND ACCRETION/OUTFLOW

- ◆ proxy accretion/outflow:
H α emission line

(Reipurth et al. 1996)



FORBIDDEN EMISSION LINES (FELs) AND OUTFLOW



σ Ori

(Rigliaco et al. 2009)

- ◆ identify outflow
- ◆ implications on H α profile
- ◆ physical properties from line ratio (n_e , \dot{M}_{wind} , ...)
- ◆ variability analysis

SELECTED OBJECTS

- ◆ GES + Chandra + COROT data
 - CSI 2264 project (P.I. G. Micela - J. Stauffer)
 - M. Guarcello
(with E. Flaccomio: Guarcello et al. in prep.)
- ◆ Spurious lines due to sky subtraction
- ◆ Herbig Haro objects (protostellar jets)

SELECTED OBJECTS

(More selected: "gir" Vs. "gar")

cNAME	Ha_class*	Ha_intr*
06411725+0954323	E	1
06410577+0948174	E	1
06405884+0930573	EE	1
06411678+0927301	EE	1
06403164+0948233	EE	1
06405159+0928445	E	1
06413656+0921514	N	NULL
06404487+1014114	EE	1
18040429-2425040	EN	1
18042795-2427599	E	1

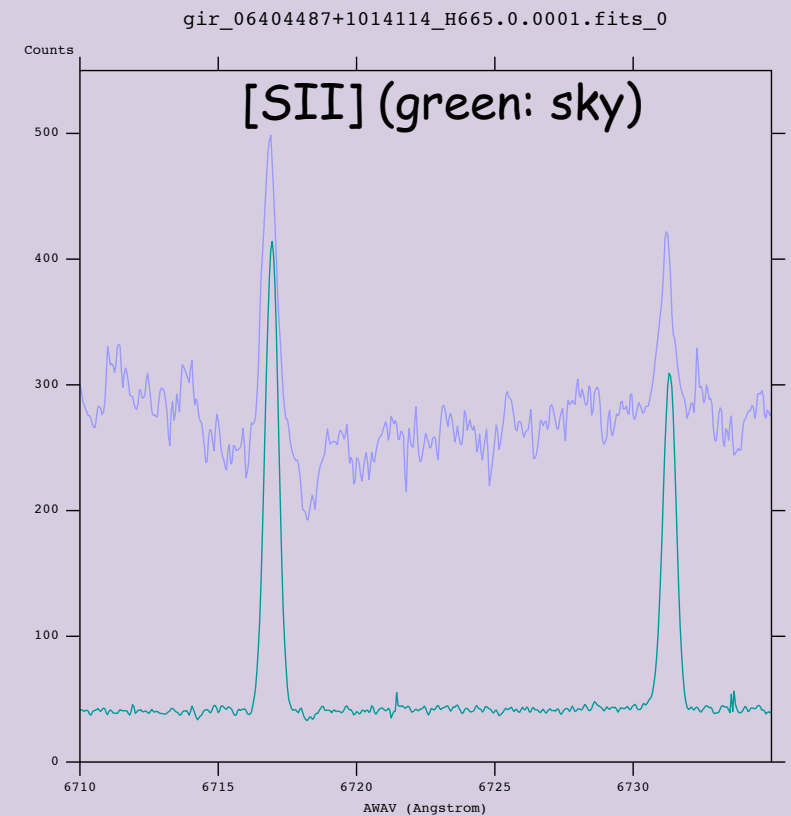
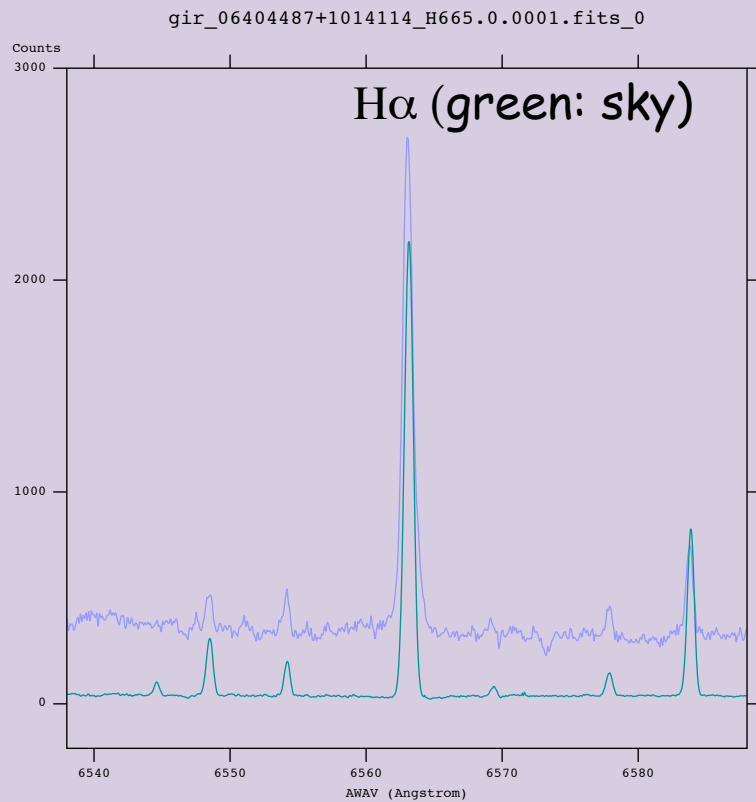
(*see also T. Zwitter's talk and Traven et al. in prep.)

SELECTED OBJECTS

cNAME	Ha_class	Ha_intr	
06411725+0954323	E	1	
06410577+0948174	E	1	
06405884+0930573	EE	1	
06411678+0927301	EE	1	X-ray
06403164+0948233	EE	1	
06405159+0928445	E	1	
06413656+0921514	N	NULL	Nebular
06404487+1014114	EE	1	HH jet
18040429-2425040	EN	1	HH jet
18042795-2427599	E	1	E

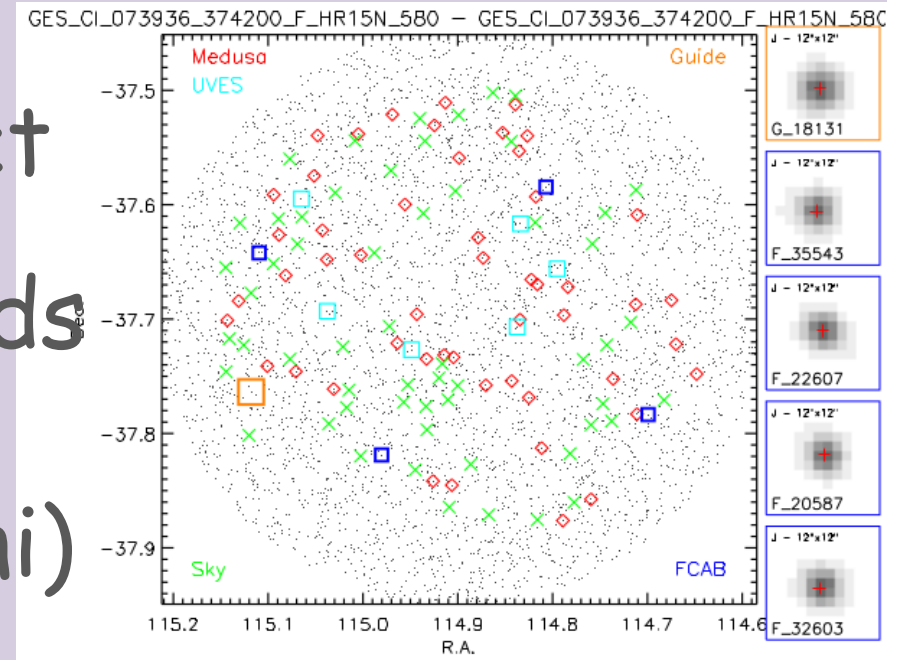
SKY SUBTRACTION & EMISSION LINES

- ◆ Nebular contribution to both $H\alpha$ and FELs
 - Strongly varying in space and time
 - Narrower than stellar ($H\alpha$)



STRATEGY

- ◆ sky fiber near each object
 - possible for sparse fields (e.g. NGC 2451 A & B
OBs Bonito & Franciosini)

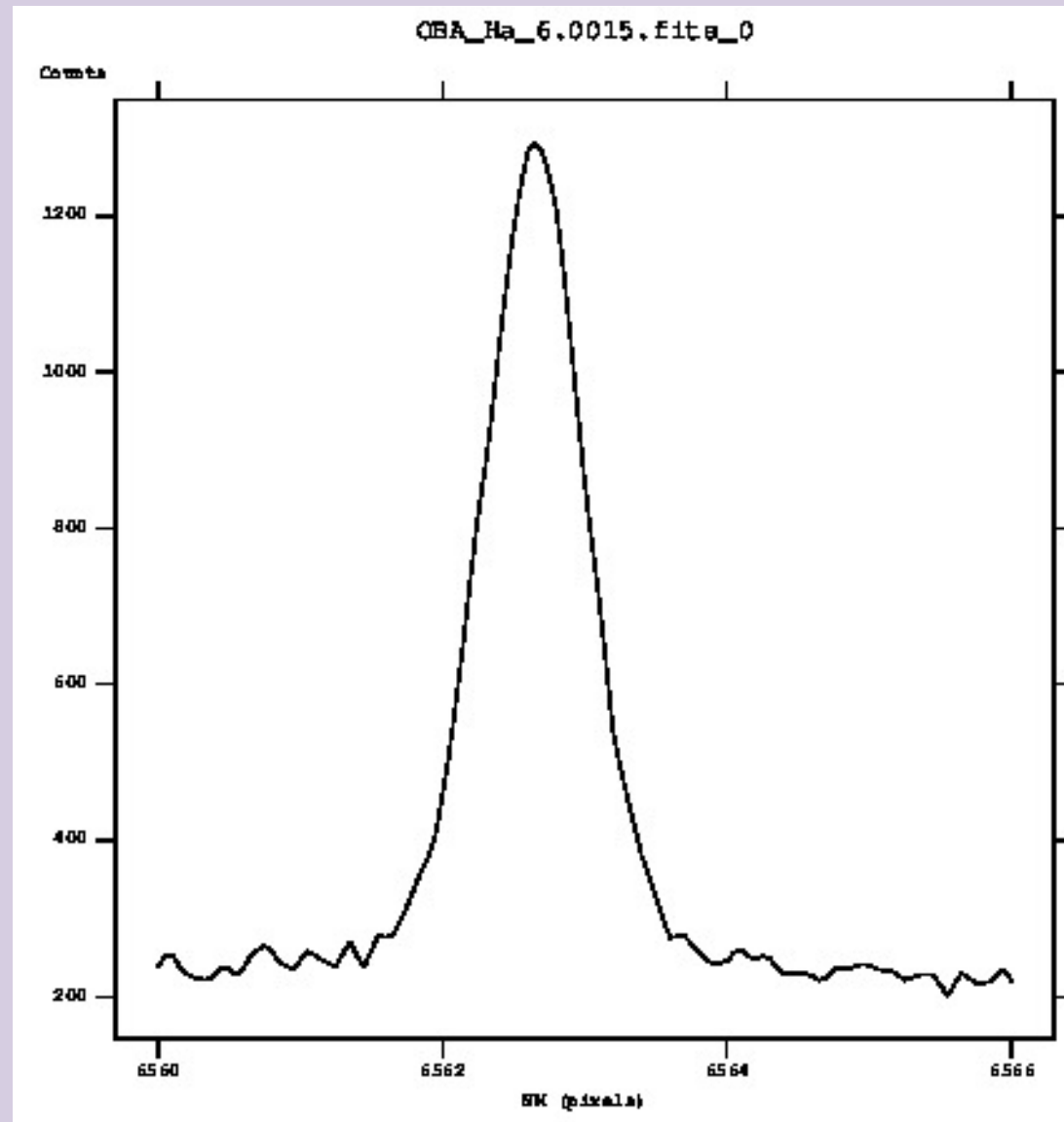


- ◆ FWZI method
(as EW and $H\alpha_{10\%}$ cannot be derived)

Automatic procedure

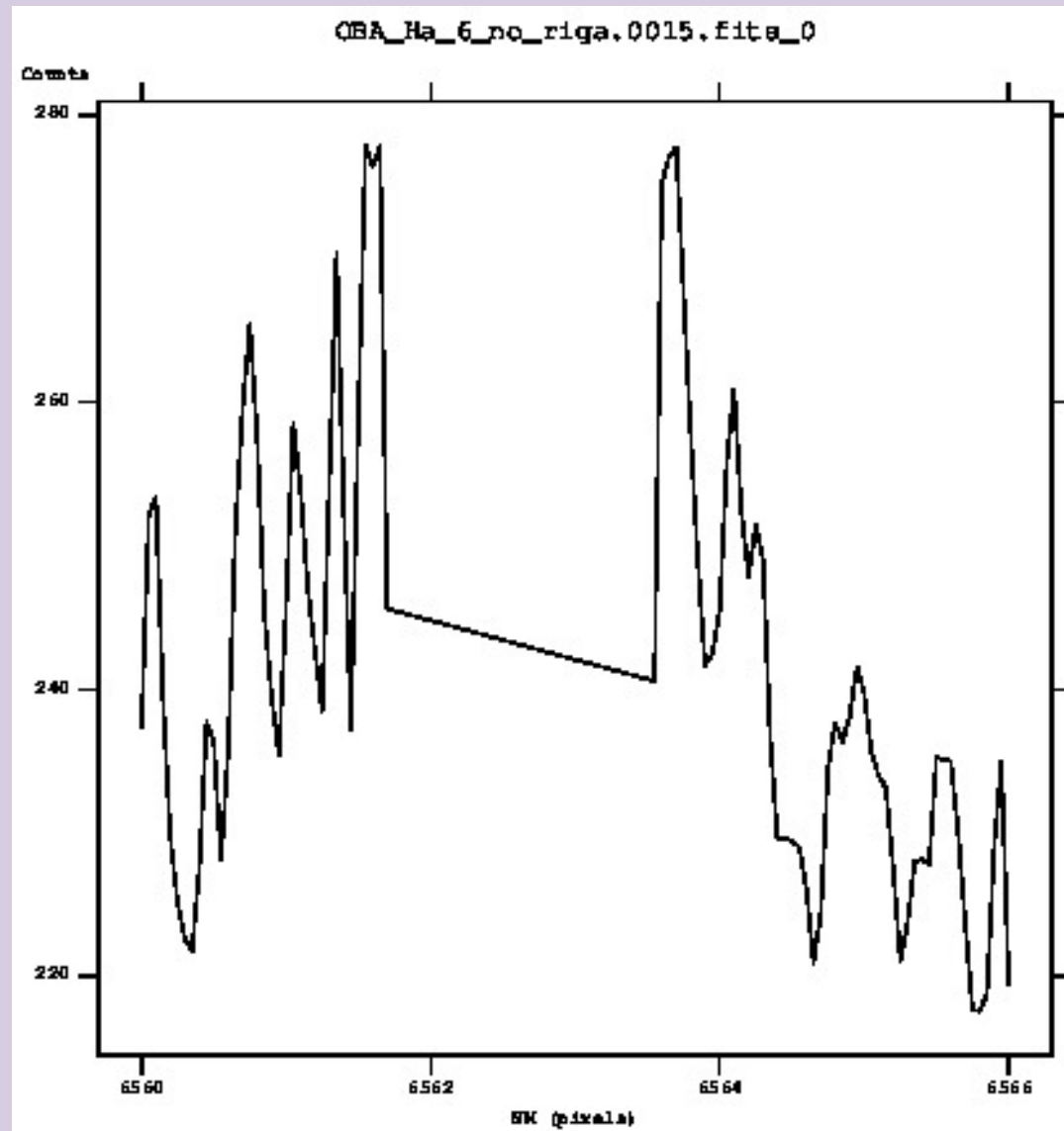
FWZI METHOD (1)

- ◆ Zoom around $H\alpha$ in sky spectra



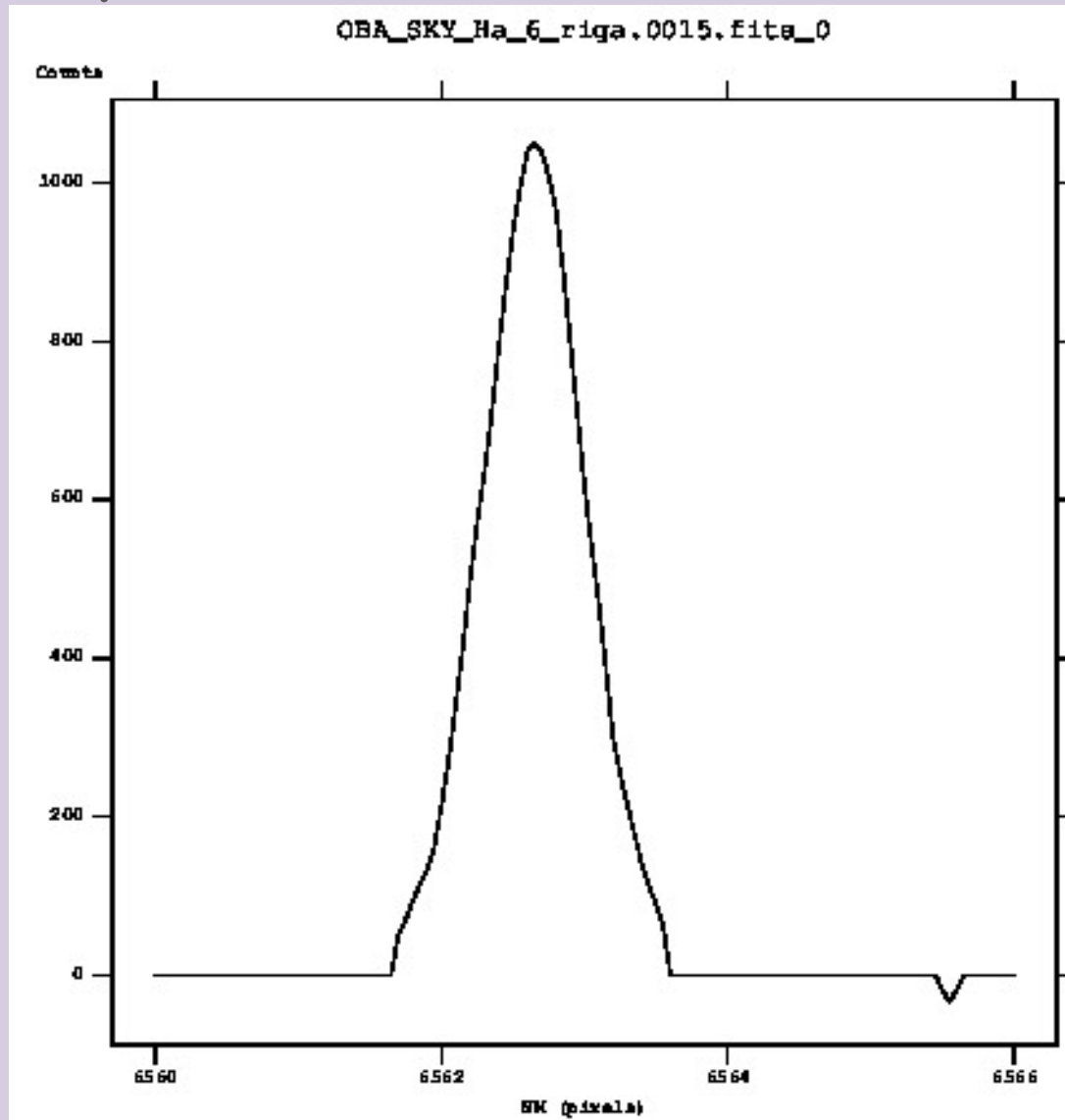
FWZI METHOD (2)

- ◆ Normalize: spectrum without the emission line



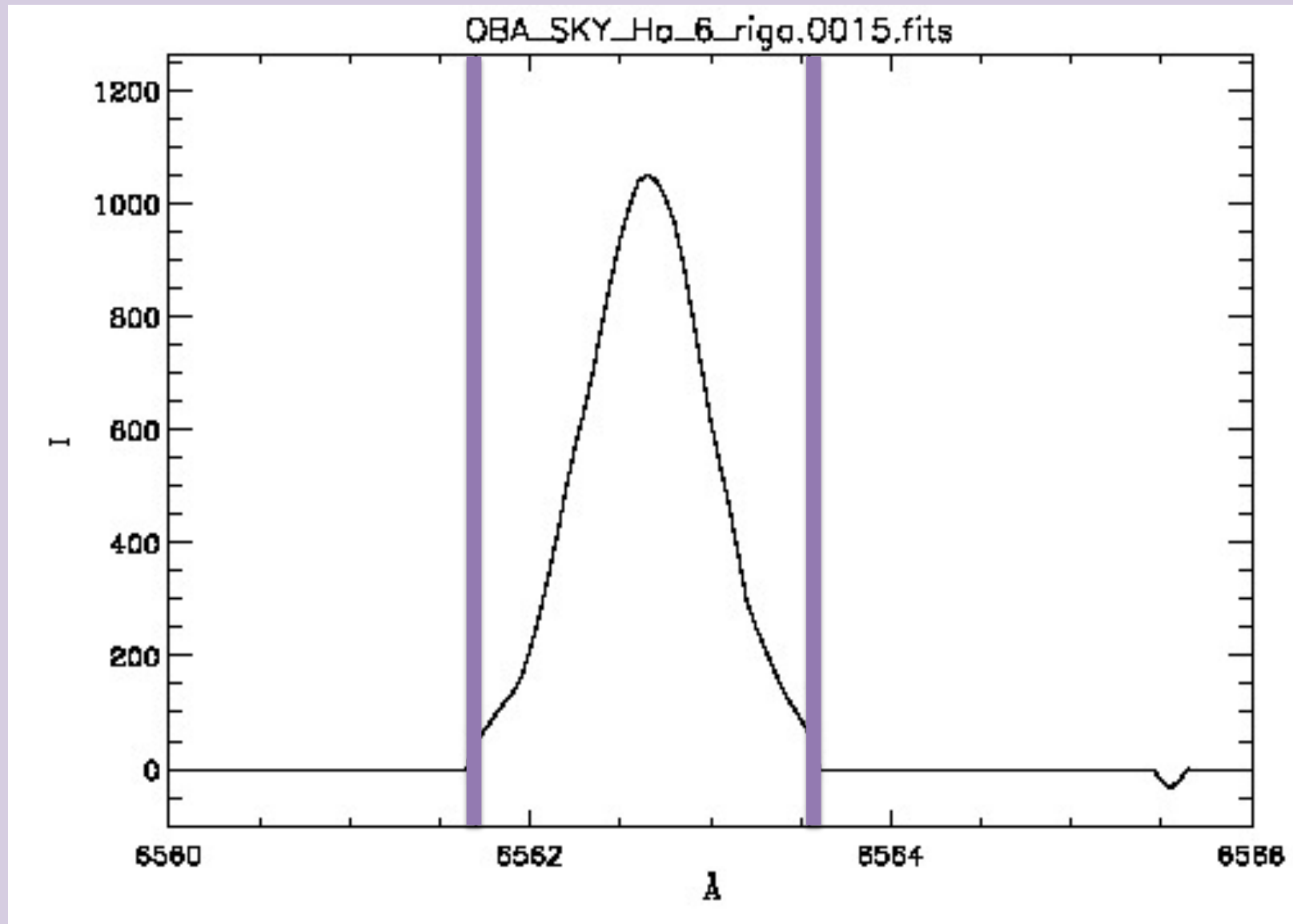
FWZI METHOD (3)

- ◆ Difference or ratio original spectrum/spectrum without the emission line



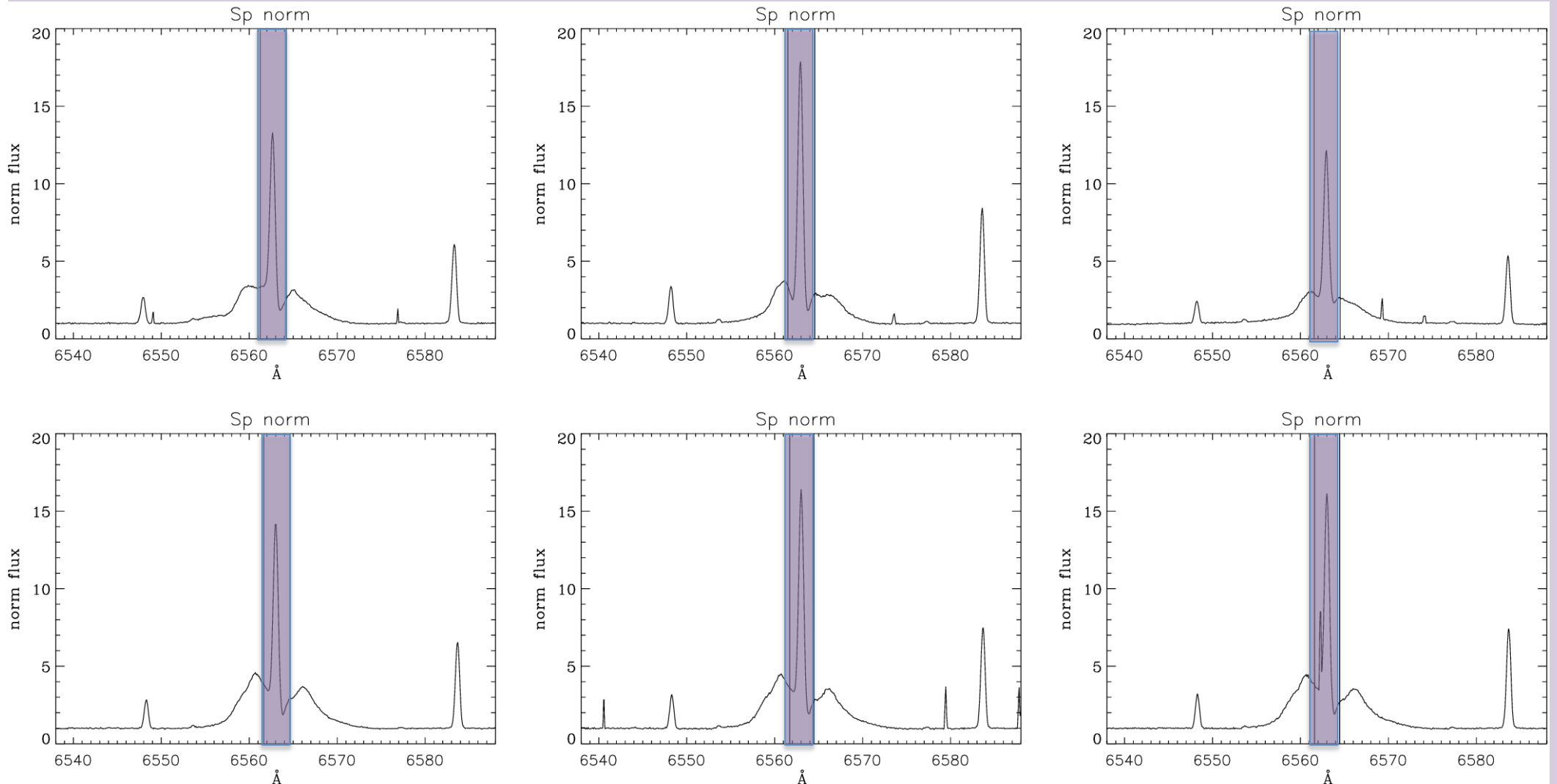
FWZI METHOD (4)

- ◆ FWZI(H α) sky < 3 Å



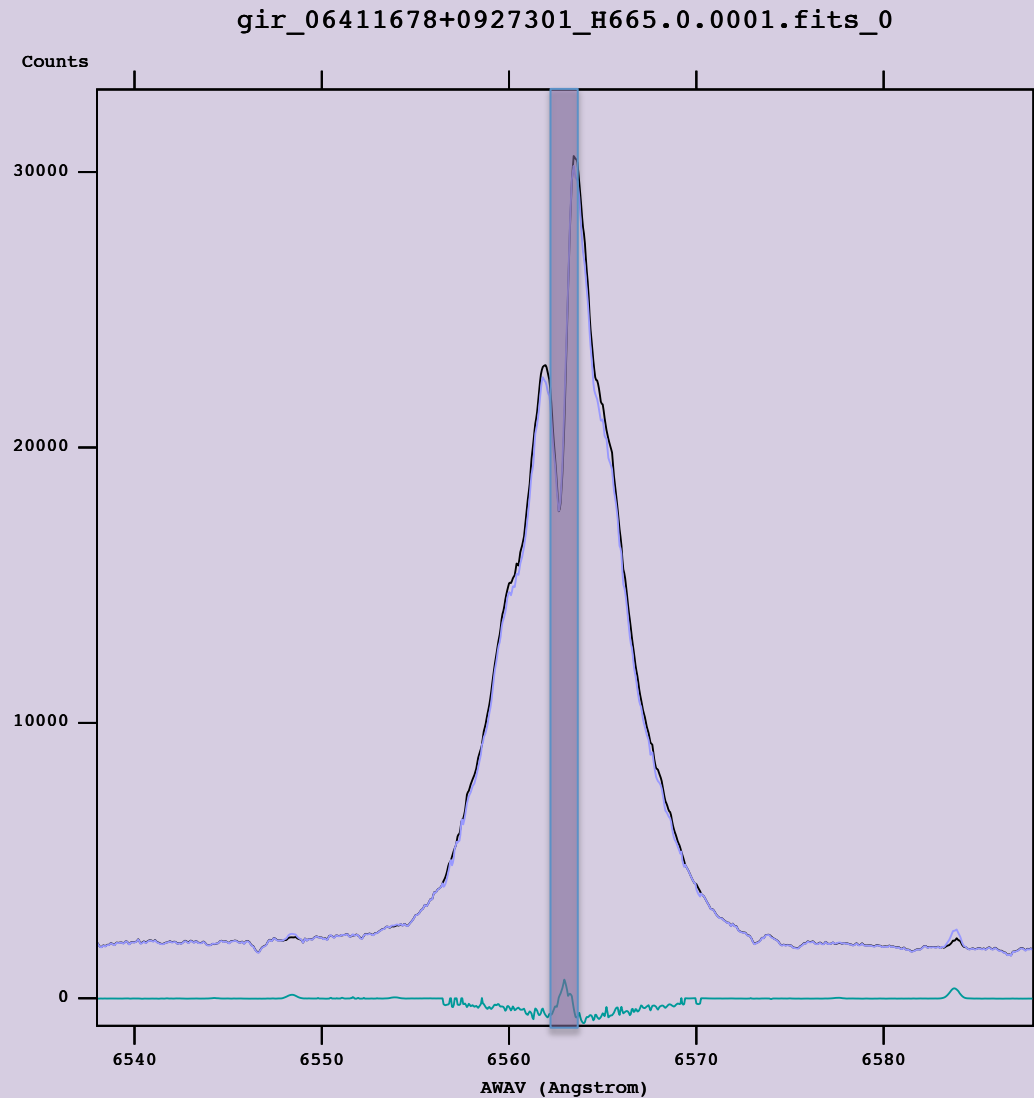
FWZI($H\alpha$) & ACCRETION

- ◆ FWZI($H\alpha$) star \gg FWZI($H\alpha$) sky
(NGC 6611, Bonito et al. 2013)



FWZI($H\alpha$) & ACCRETION

- ◆ FWZI($H\alpha$) star \gg FWZI($H\alpha$) sky



NGC 2264:
X-ray

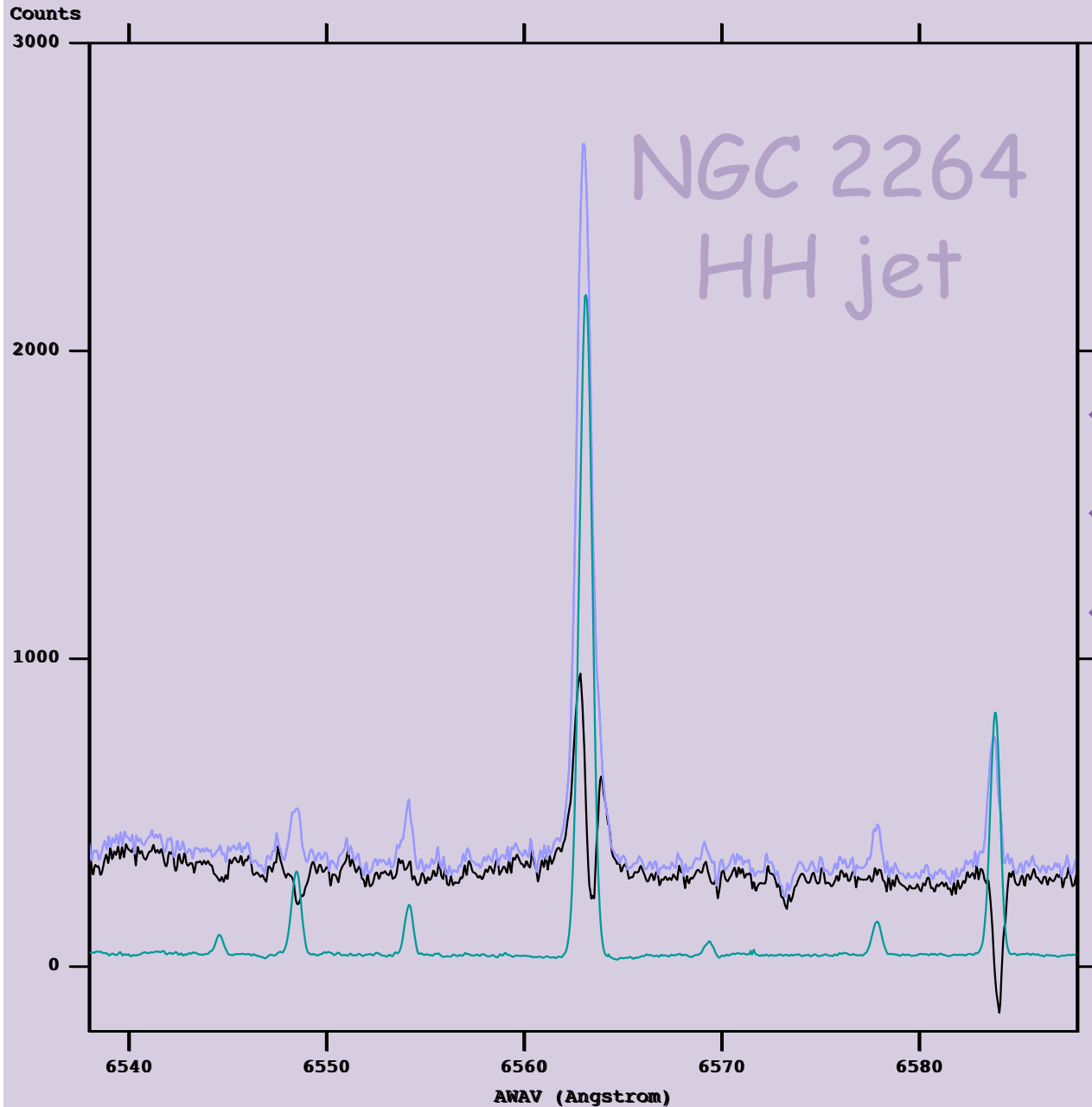
(Bonito et al. in prep.)

SKY SUBTRACTION: SPURIOUS LINES

- ◆ H α absorption line:
 - ✓ implication on accretion/outflow activity
- ◆ [SII] and [NII] absorption lines:
 - ✓ implication on outflow activity
- ◆ Ca I absorption line near Li:
 - ✓ implication on age and membership

SKY SUBTRACTION AND H α + [NII]

gir_06404487+1014114_H665.0.0001.fits_0



- ◆ GES spectrum
- ◆ Original spectrum
- ◆ SKY spectrum

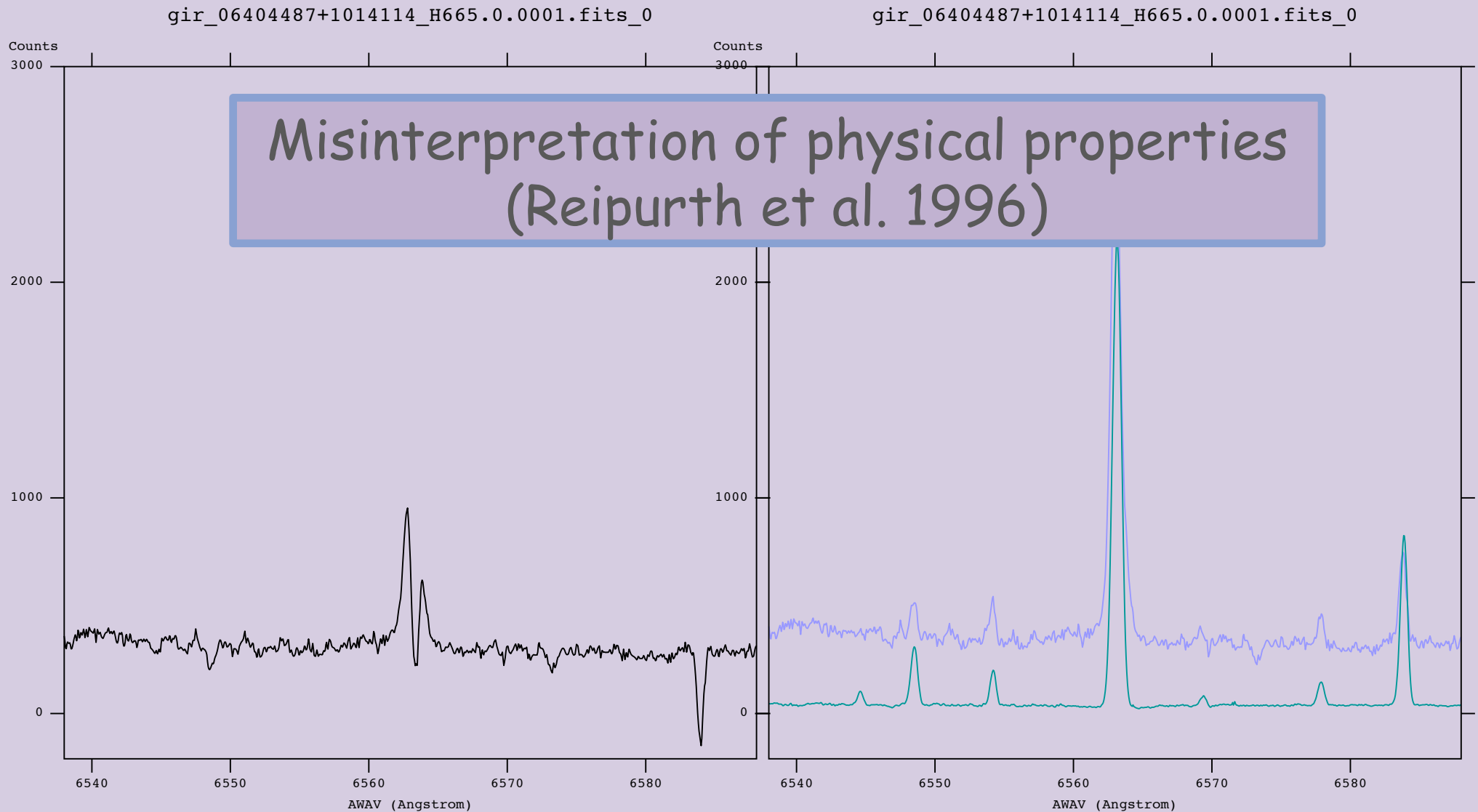
(Bonito et al. in prep.)

SKY SUBTRACTION AND H α + [NII]

◆ GES spectrum

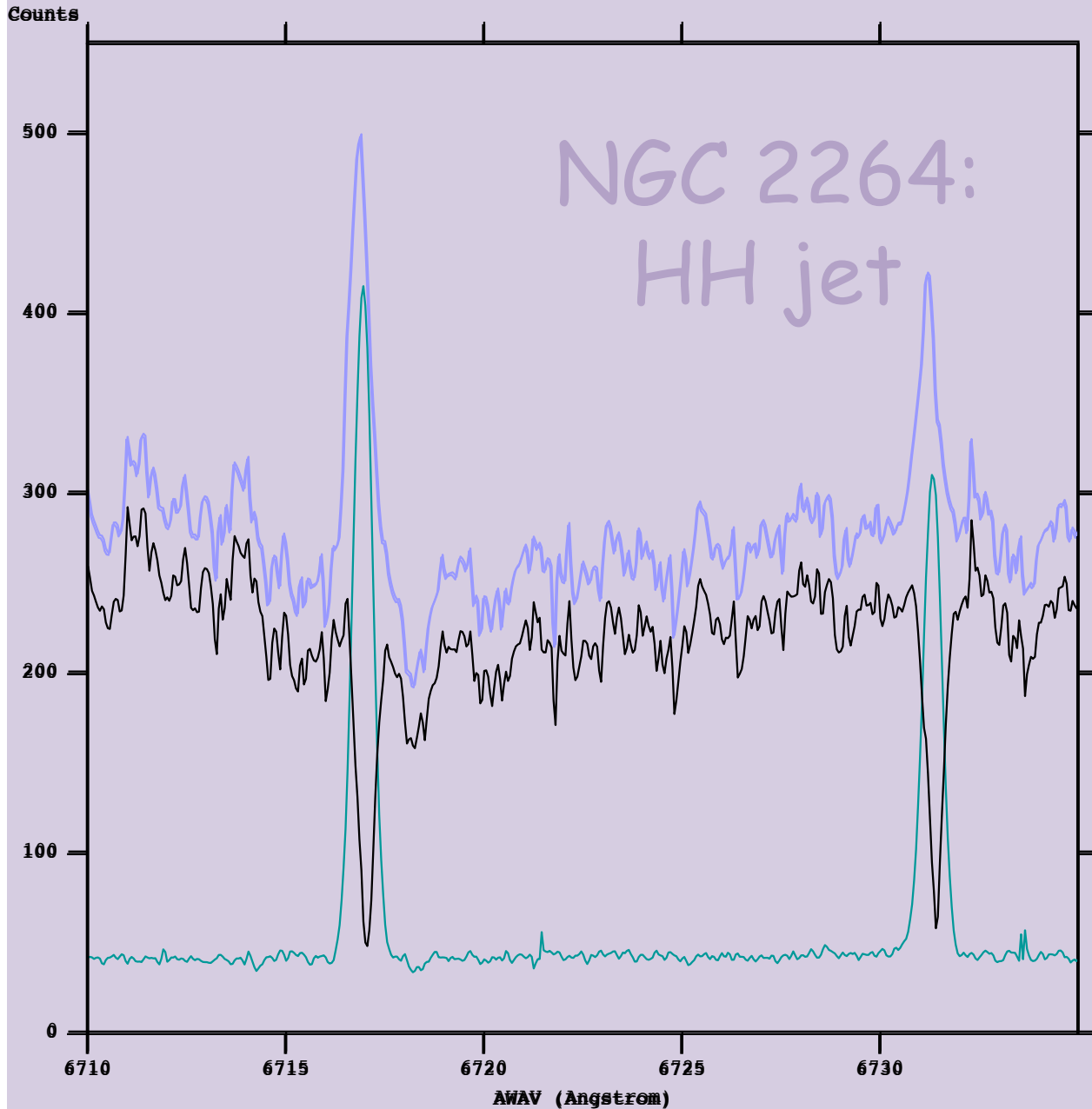
◆ Original spectrum

◆ SKY spectrum



SKY SUBTRACTION AND [SII]

gir_06404487+1014114_H665.0.0001.fits_0



- ◆ GES spectrum
- ◆ Original spectrum
- ◆ SKY spectrum

(Bonito et al. in prep.)

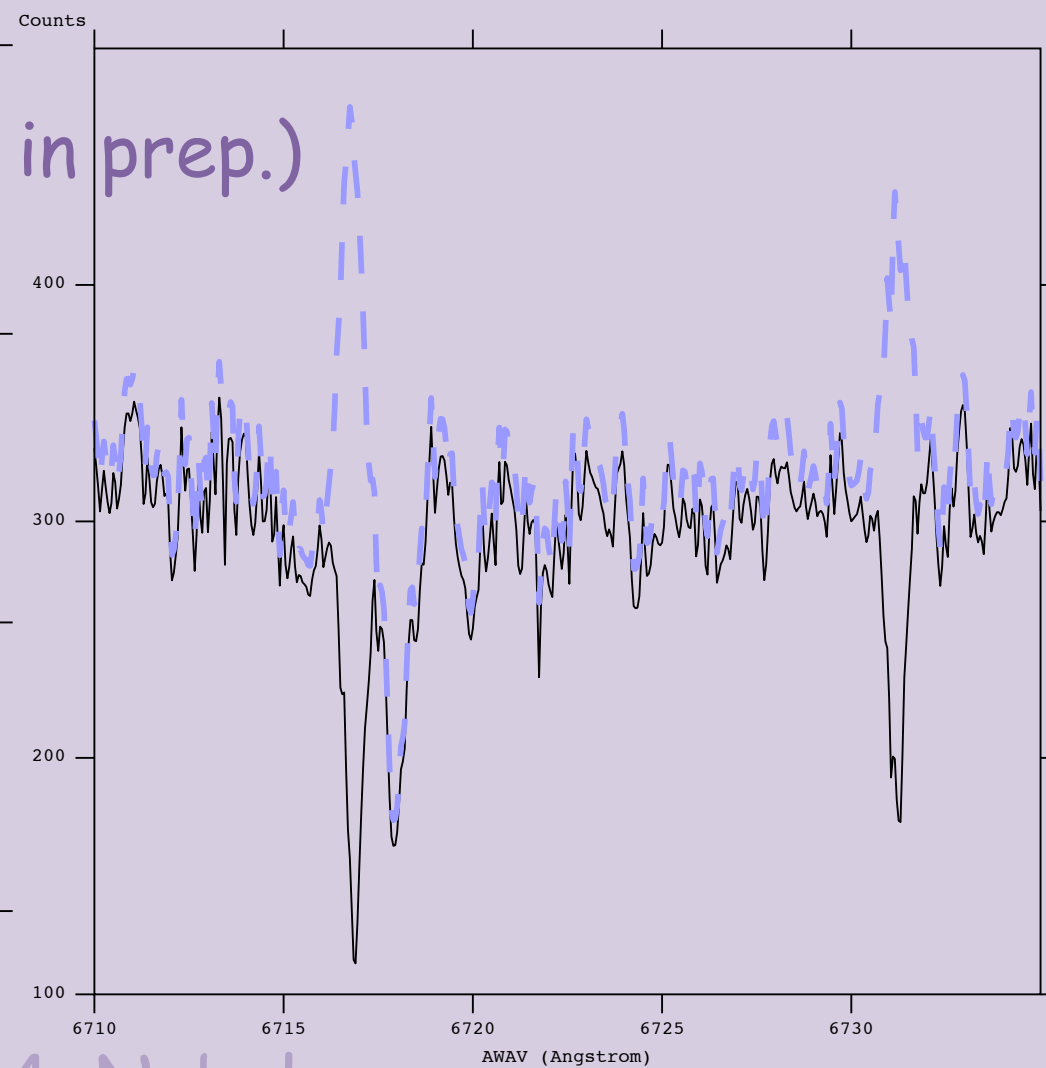
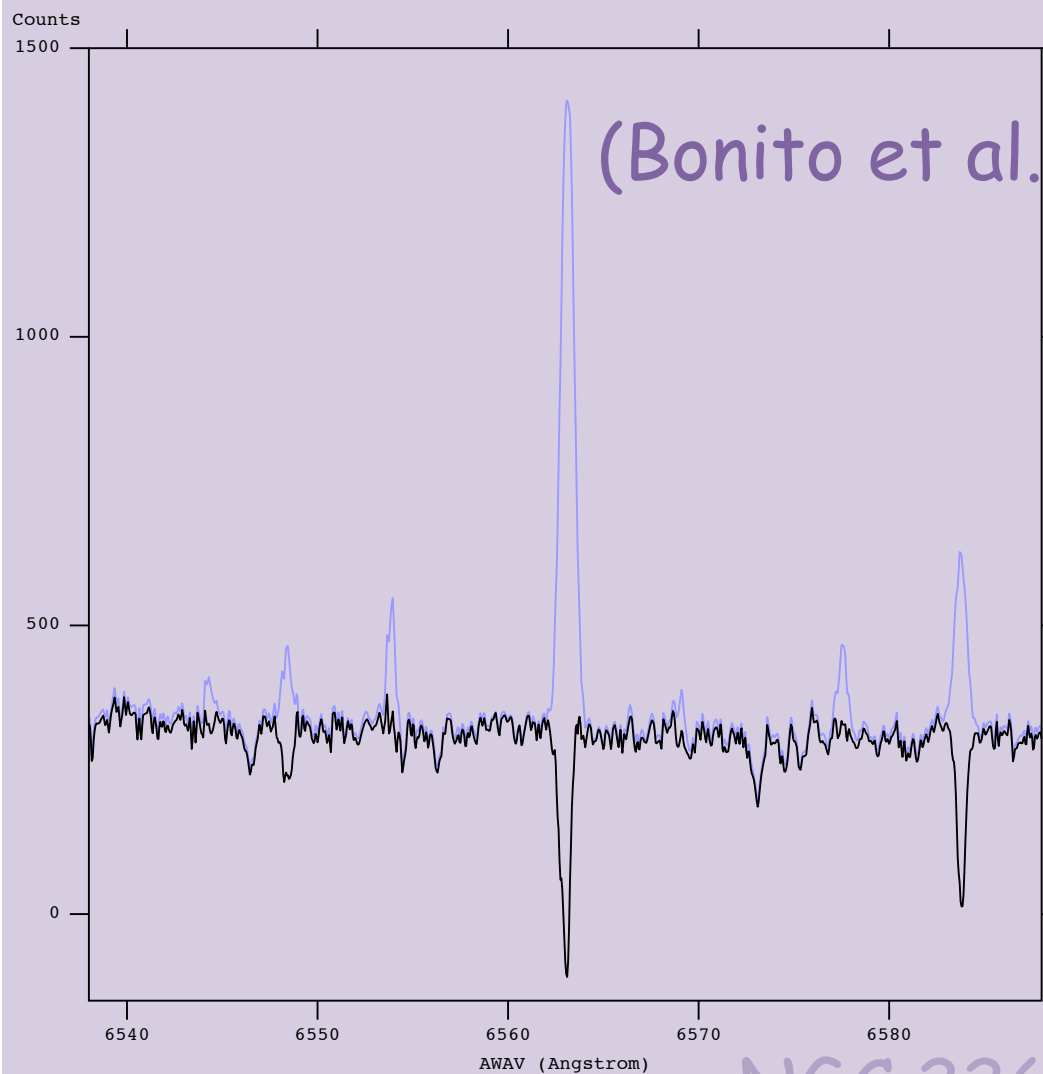
SKY SUBTRACTION

◆ GES spectrum

◆ Original spectrum

C20121205_00020_st.0029.fits_0

gir_06413656+0921514_H665.0.0001.fits_0



NGC 2264: Nebular
(Ha_class = N, Traven et al. in prep.)

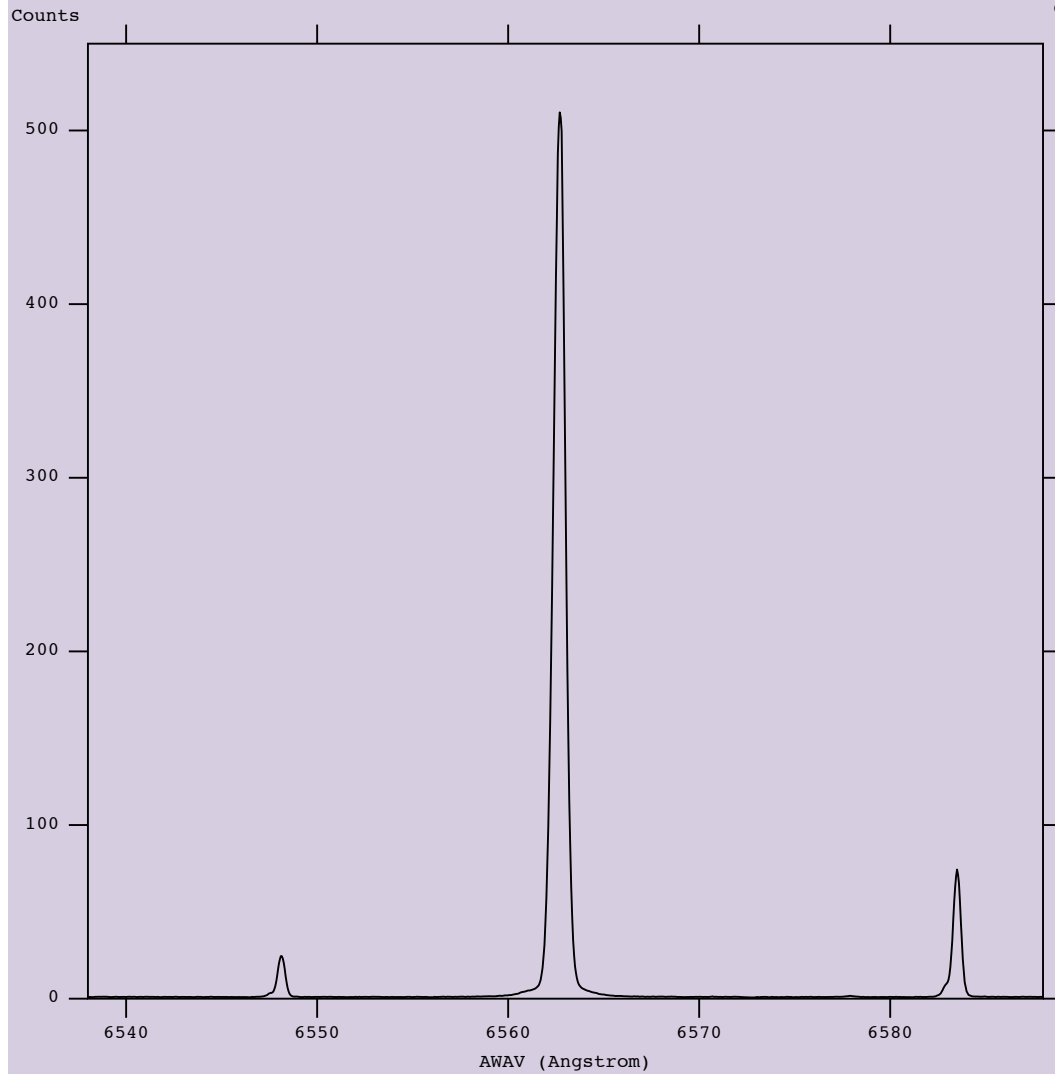
RESULTS

(Bonito et al. in prep.)

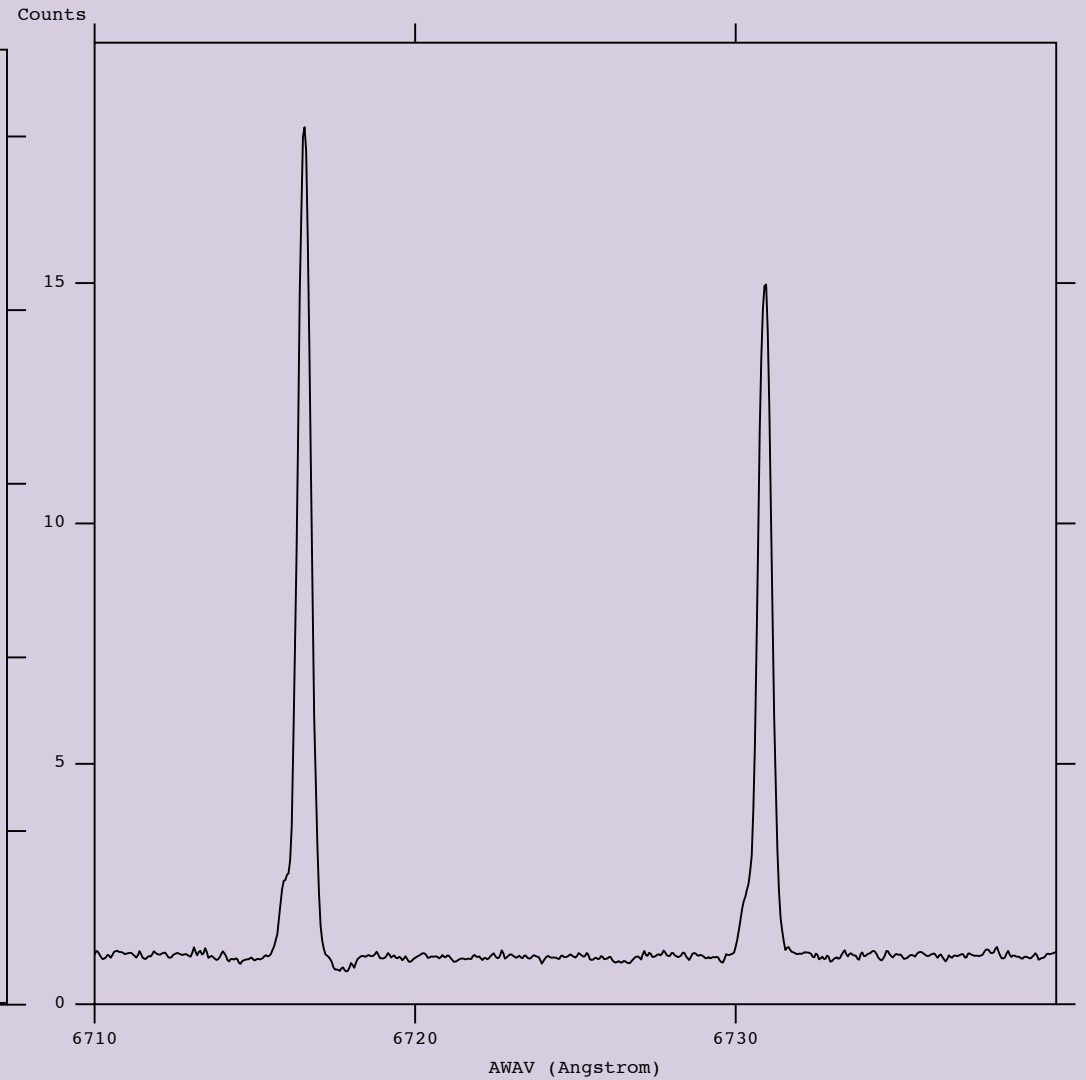
NGC 6530

HH jet, EN

gir_18040429-2425040_H665.0_norm.0001.fits_0



gir_18040429-2425040_H665.0_norm.0001.fits_0



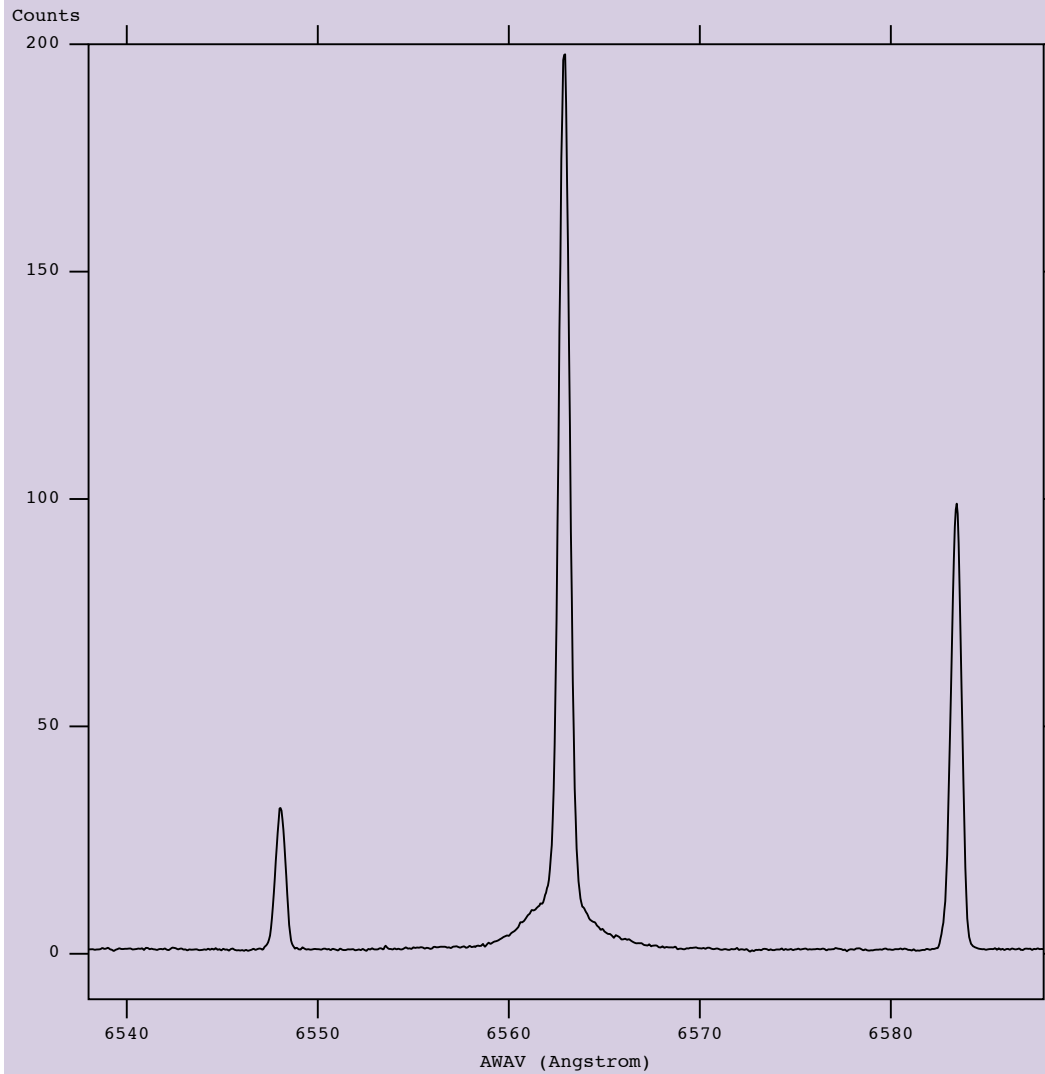
RESULTS

(Bonito et al. in prep.)

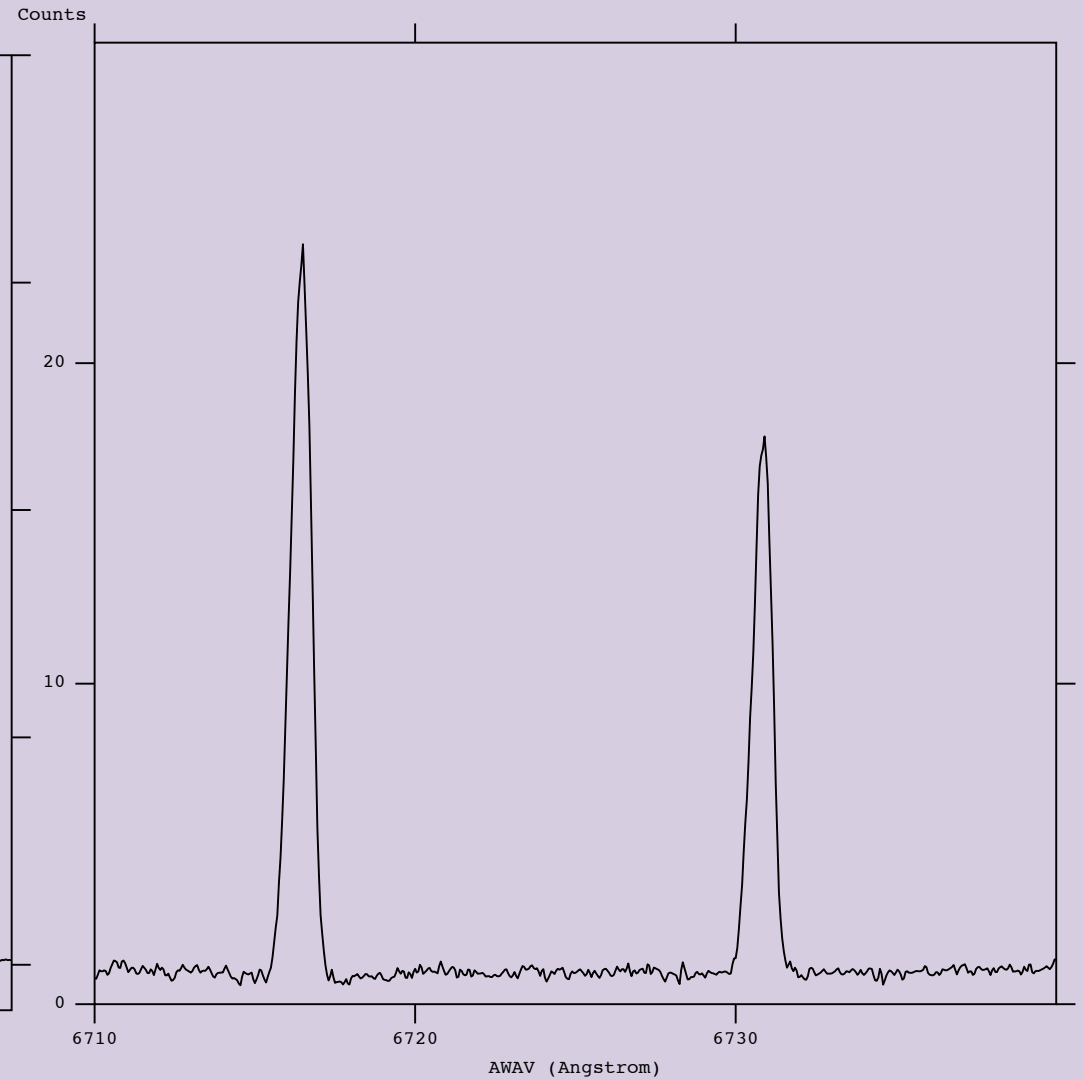
NGC 6530

E

gir_18042795-2427599_H665.0_norm.0001.fits_0



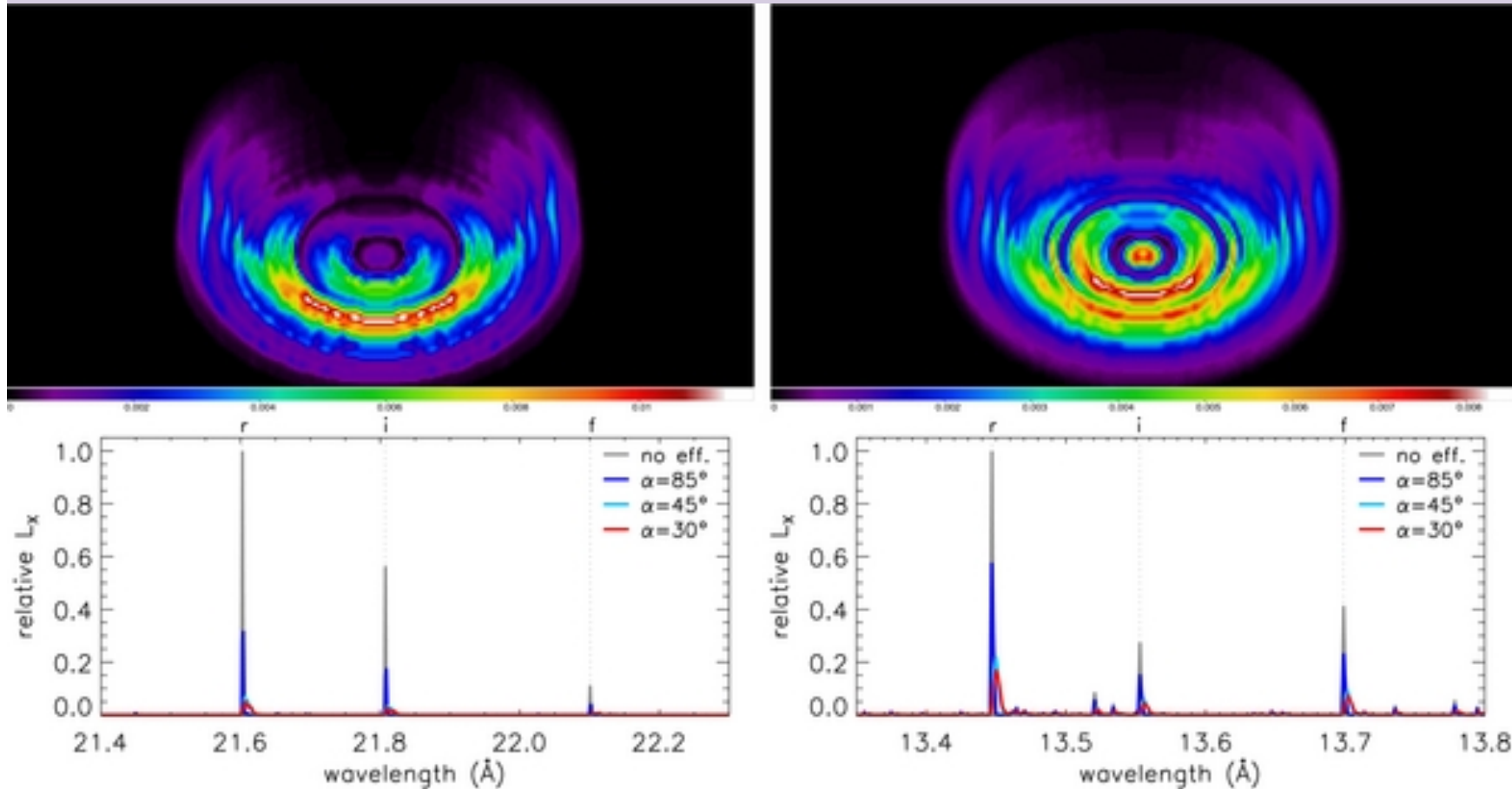
gir_18042795-2427599_H665.0_norm.0001.fits_0



CONCLUSIONS

- ◆ **Accretion/outflow** in GES young clusters
- ◆ **Sky subtraction**
 - **FWZI**(H α) method
 - **Spurious** absorption lines:
FELs useful both for physics
(outflow activity) and as a proxy
for sky subtraction:
flag if absorption:
warning better to use the original spectra

FUTURE PERSPECTIVES



(Bonito et al. 2014, ApJL)

Laboratory formation of a scaled protostellar jet by coaligned poloidal magnetic field

[B. Albertazzi](#), [A. Ciardi](#), [M. Nakatsutsumi](#), [T. Vinci](#), [J. Béard](#), [R. Bonito](#), [J. Billette](#) et al.

Science 2014

FUTURE PERSPECTIVES

- ◆ Looking for outflows
- ◆ FWZI($H\alpha$) method for accretion

please contact me
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