Changing Phases of Kepler Exoplanets: Evidence for Clouds and Winds

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Using 18 guarters of data from the Kepler Space Telescope we analyze the transit and phase curve of 14 systems with significant detections in each of the phase curve components: the planet's phase function, Doppler boosting and ellipsoidal variations. We model the full phase curve simultaneously, including primary and secondary transits, and derive albedos, dayand night-side temperatures, planet masses and planetary light peak offsets.





Eastward Offset for Hotter Planets

For tidally locked planets, an eastward shift in the substellar hot-spot is predicted by circulation models (e.g Cooper et al. 2005; Heng et al. 2011).

 Could be caused by winds moving in the direction of planetary rotation.

For more information:

Esteves, De Mooij and Jayawardhana. arXiv:1407.2245, (2014)

References

Cooper et al. 2005, ApJ, 629, L45 Demory et al. 2013, ApJ, 776, L25 Heng et al. 2011, MNRAS, 413, 2380



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Calculated reflected light fraction is marginally consistent with:

- Cooler Planets Mostly Reflected Light Clouds
- Hotter Planets Mostly Thermal **Emission Hot-Spot**



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Westward Offset for Hotter Planets Possibly caused by clouds on the morning-side (as seen for Kepler-7b;

- Demory et al. 2013).
- · Could have condensed on the nightside, when temperatures are cooler.
- Thickest at sunrise then gradually disperse throughout the day as the atmosphere heats up.

